



Particle-In-Cell Simulations of the Rippled Low Mach Number Shock in High Beta Cosmic Plasma

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KU KDM'19, Zakopane, March 6 – 8, 2019

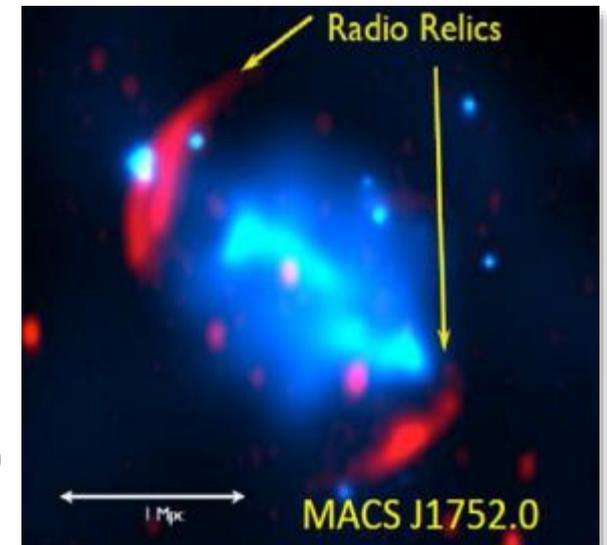
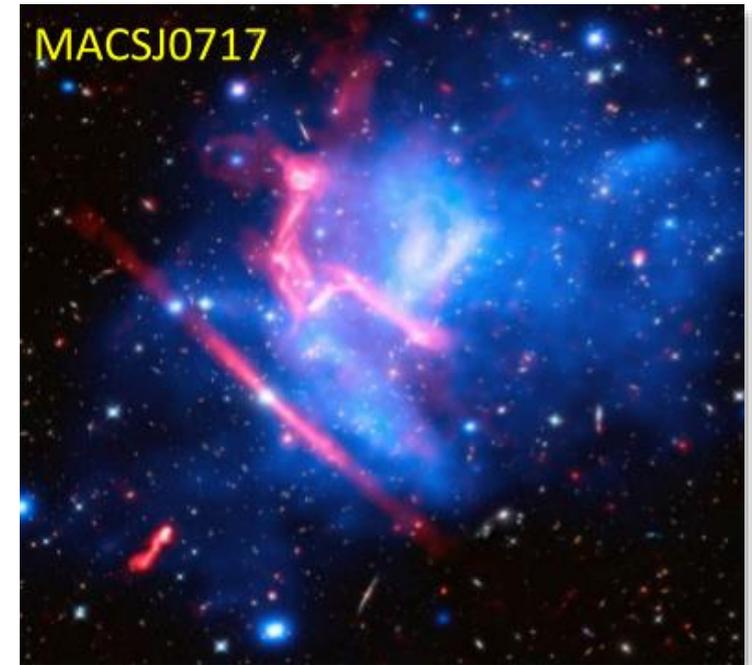
Astrophysical objects

Collisionless shocks in astrophysical objects on the various scales:

- Earth's bow shock.
- Solar wind termination shock.
- Supernova remnant (SNR) shocks.
- Active galactic nuclei (AGN) shocks.
- Large-scale structure formation shocks, mostly in the clusters of galaxies:
 - turbulence shocks,
 - infall shocks,
 - merger shocks.

In the latter case, low-Mach-number ($M \ll 10$) shocks are found to propagate in high-beta ($\beta > 1$) plasmas.

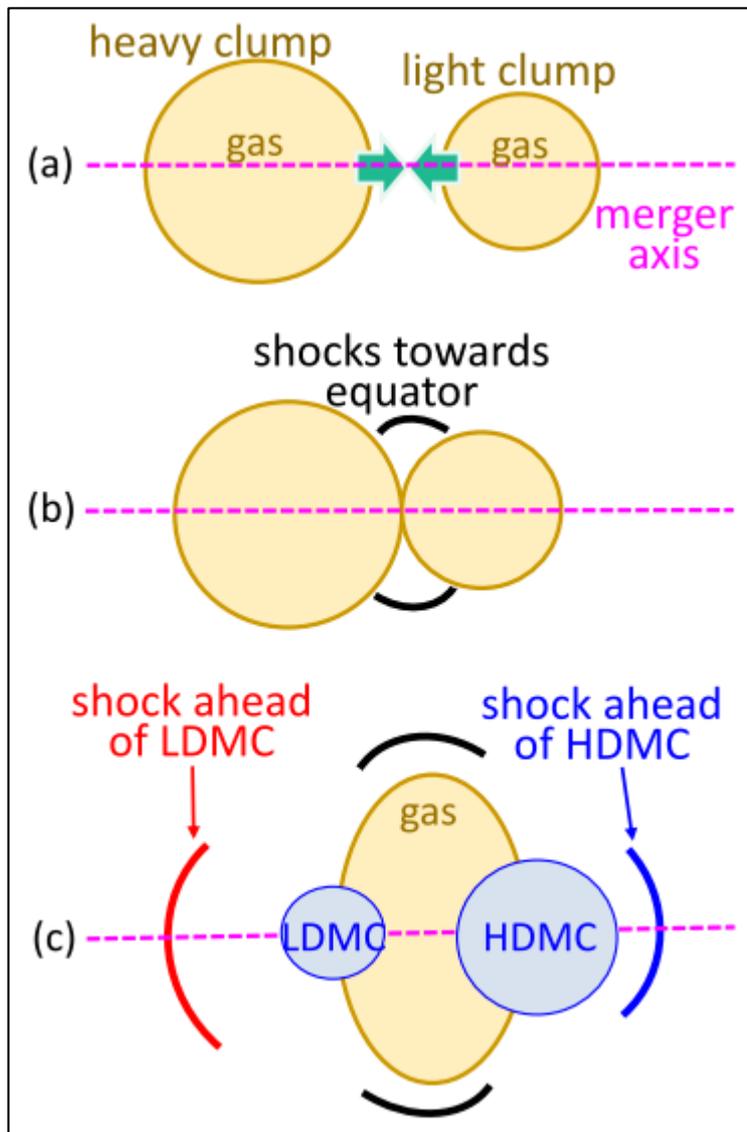
X-ray and radio emission show the electron acceleration to non-thermal energies.



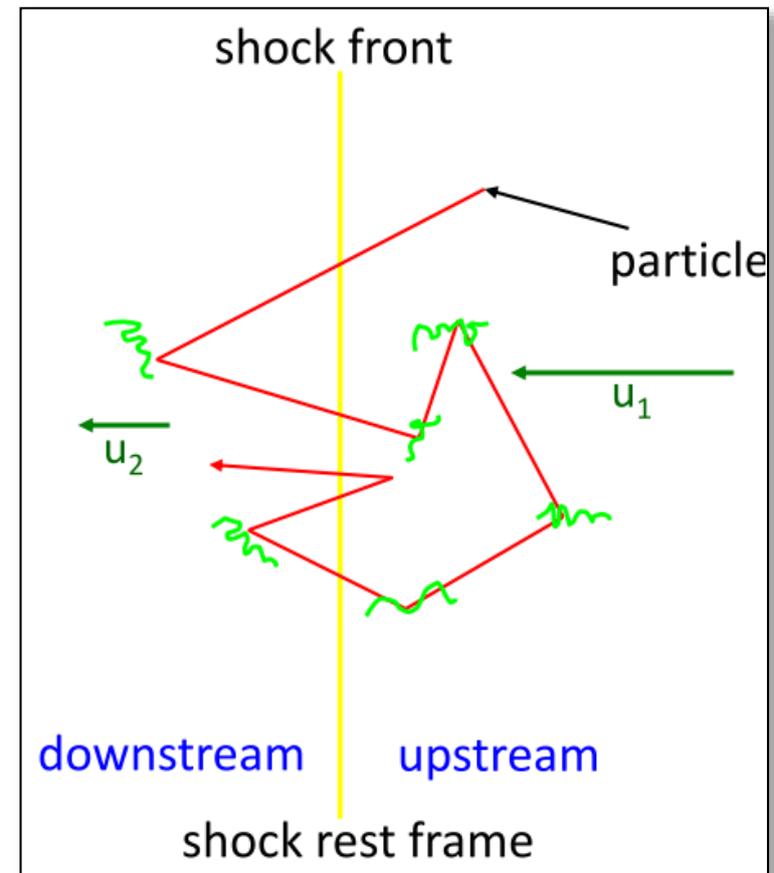
White – optical (Hubble)
Blue – X-ray (Chandra)
Red – radio (VLA)

Shock forming and particles acceleration

Simple binary galaxy merger



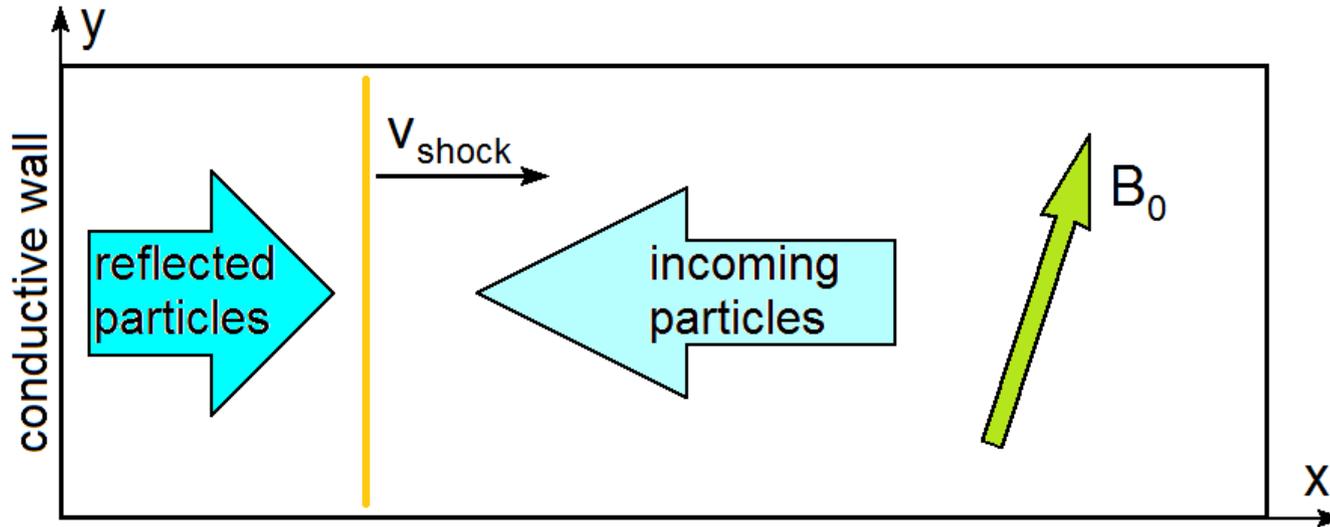
Simple scheme of the DSA



Injection problem:

Particle should be pre-accelerated to be involved in the DSA process

2D-3V Particle-In-Cell simulation



$$L_x = 65,000 \Delta \approx 433 \lambda_{si} \quad L_y = 4,800 \Delta \approx 32 \lambda_{si}$$

2-component
proton-electron
plasma.

Shock is formed via
interaction between
reflected and
incoming particles.

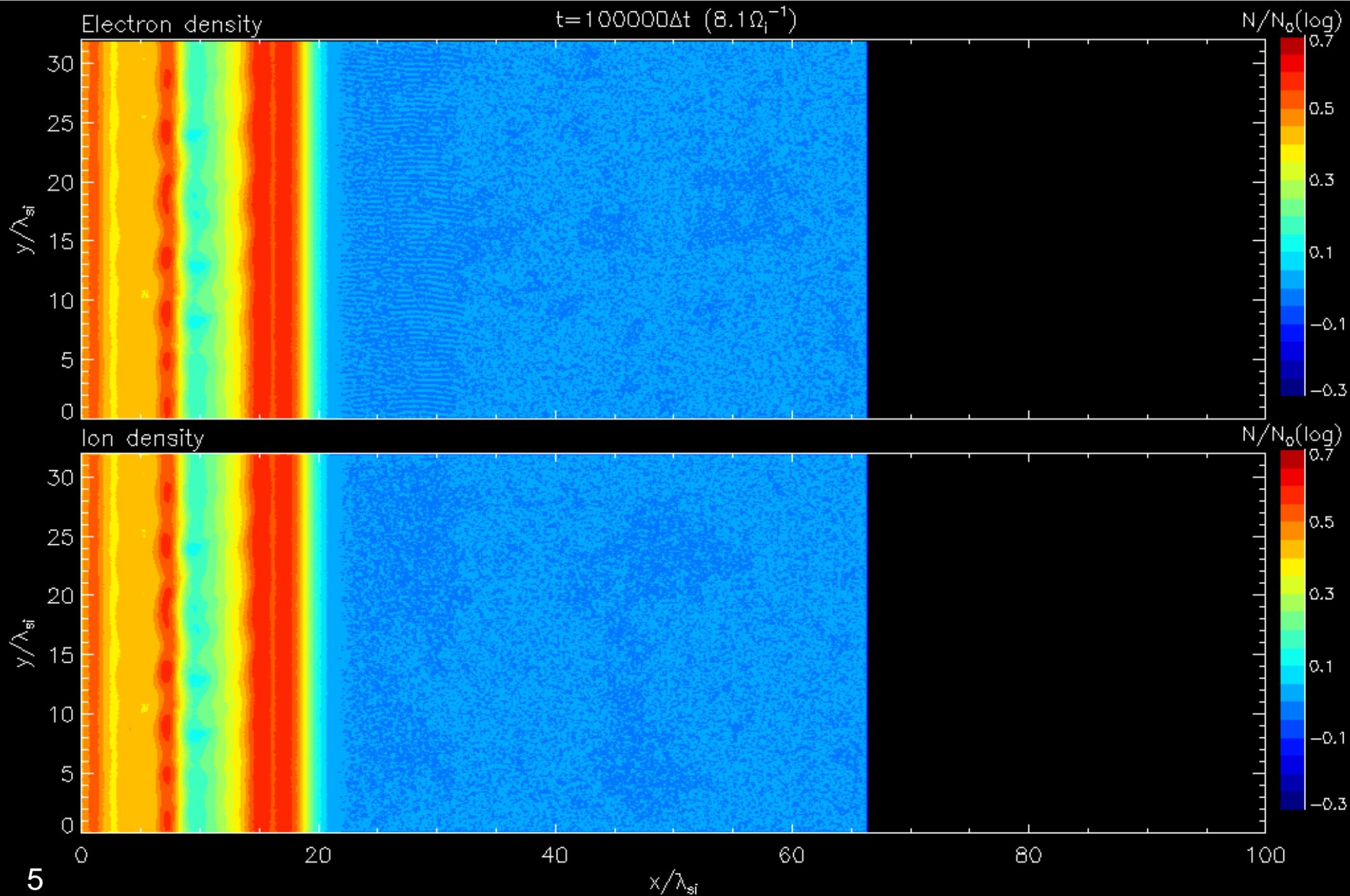
Physical parameters:

- Reduced ion to electron mass ratio $m_i/m_e = 100$
- Upstream plasma velocity $v_0 = 0.1c$
- Electron/ion thermal velocity $v_{e/i\ th} \approx 0.387c / 0.0387c$
(plasma temperature $k_B T \approx 40$ keV)
- Sonic Mach number of the shock $M_s = 3$
- Alfvén Mach number of the shock $M_A \approx 6$
- Plasma beta ($\beta = \rho_{\text{therm}} / \rho_{\text{mag}}$) $\beta = 5$
- Magnetic field orientation $\theta = 75^\circ$

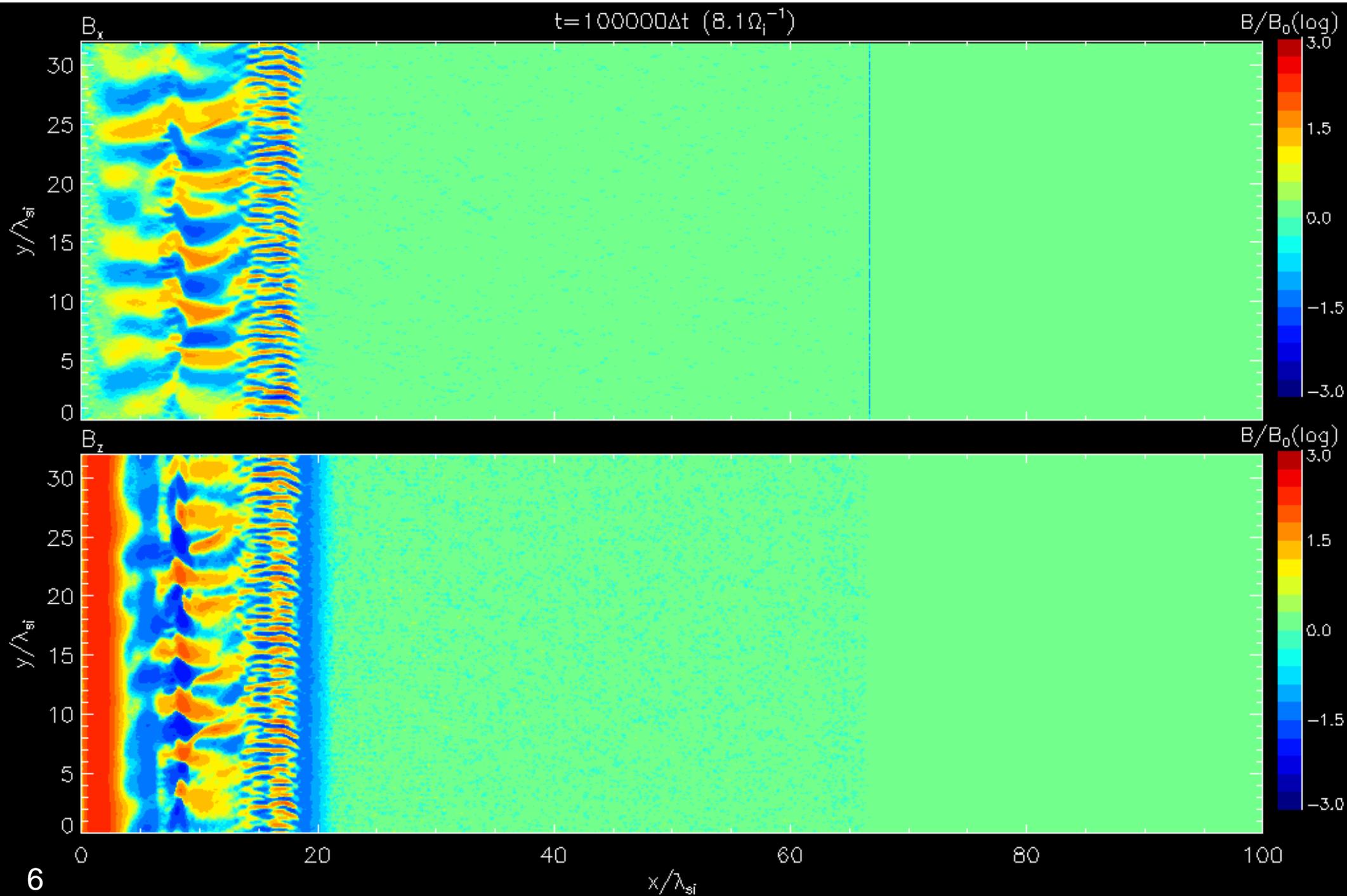
Computations:

- on *PROMETHEUS* cluster
- up to 6240 CPU cores
- ~ 10 mln of CPU-hours
- ~ 60 TB of disk space used for data output

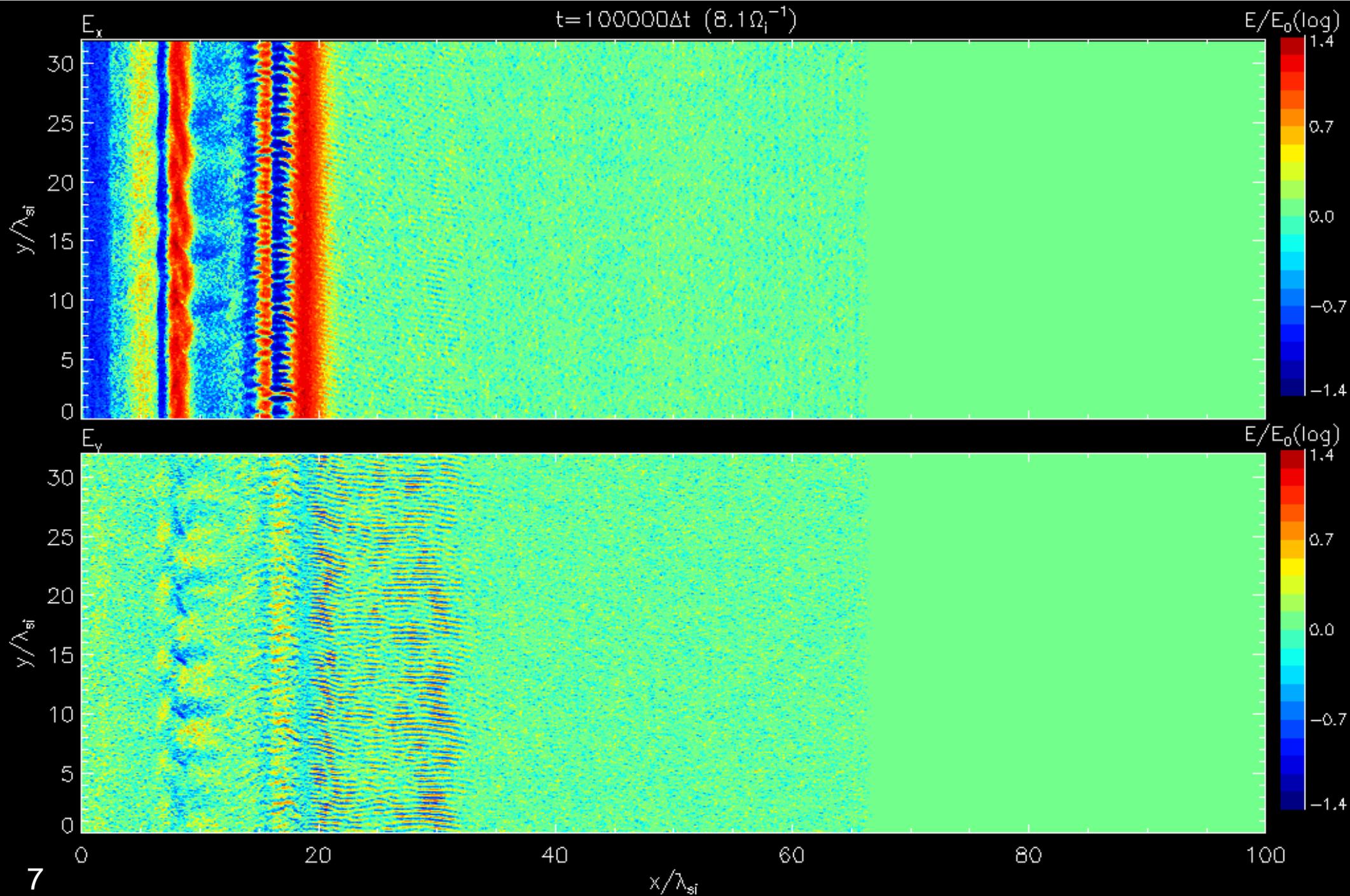
Global system evolution: *Plasma density*



Global system evolution: *Magnetic field*

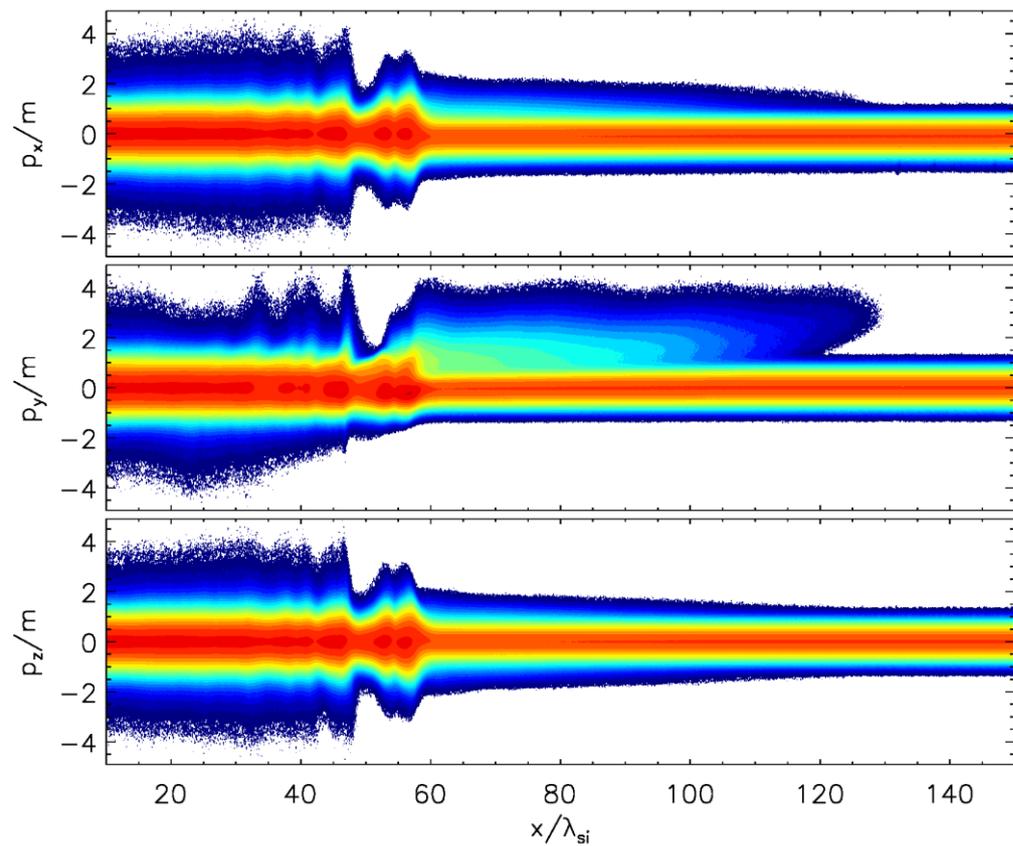


Global system evolution: *Electric field*

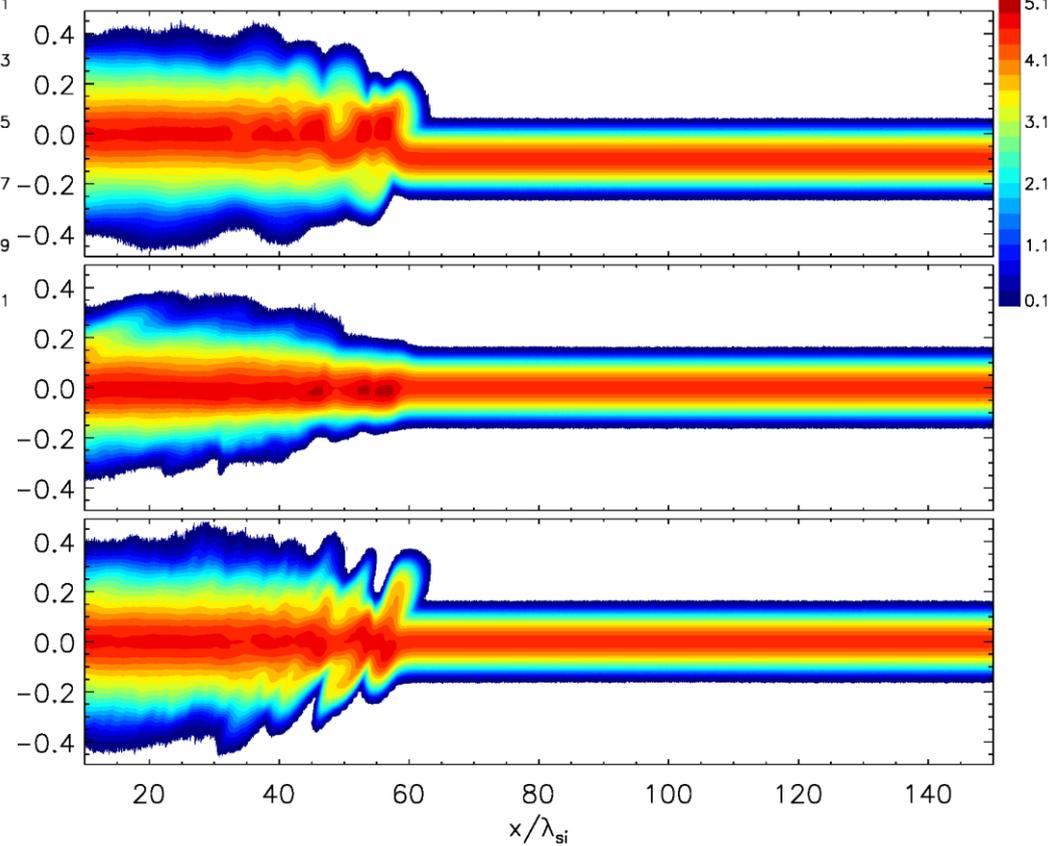


Phase space evolution: *Laminar shock*

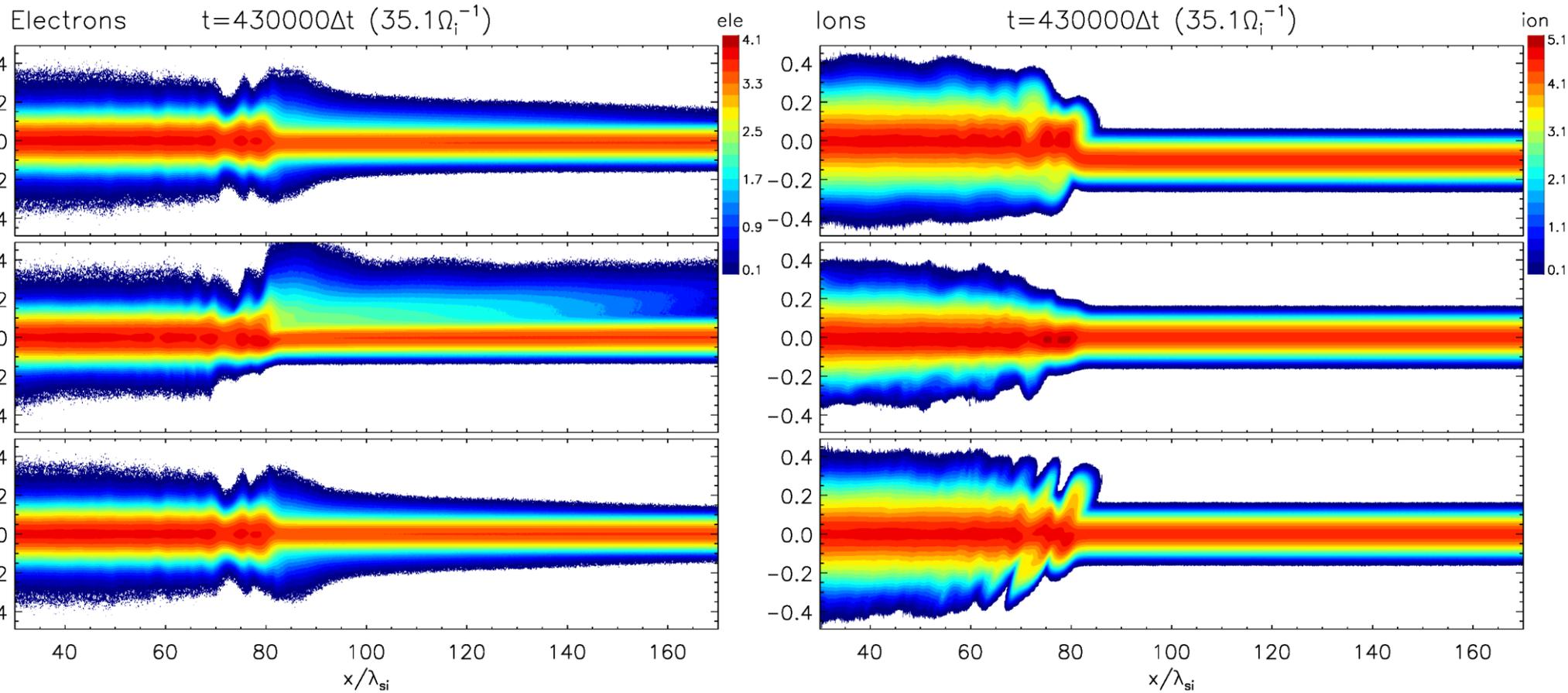
Electrons $t=300000\Delta t$ ($24.4\Omega_i^{-1}$)



Ions $t=300000\Delta t$ ($24.4\Omega_i^{-1}$)

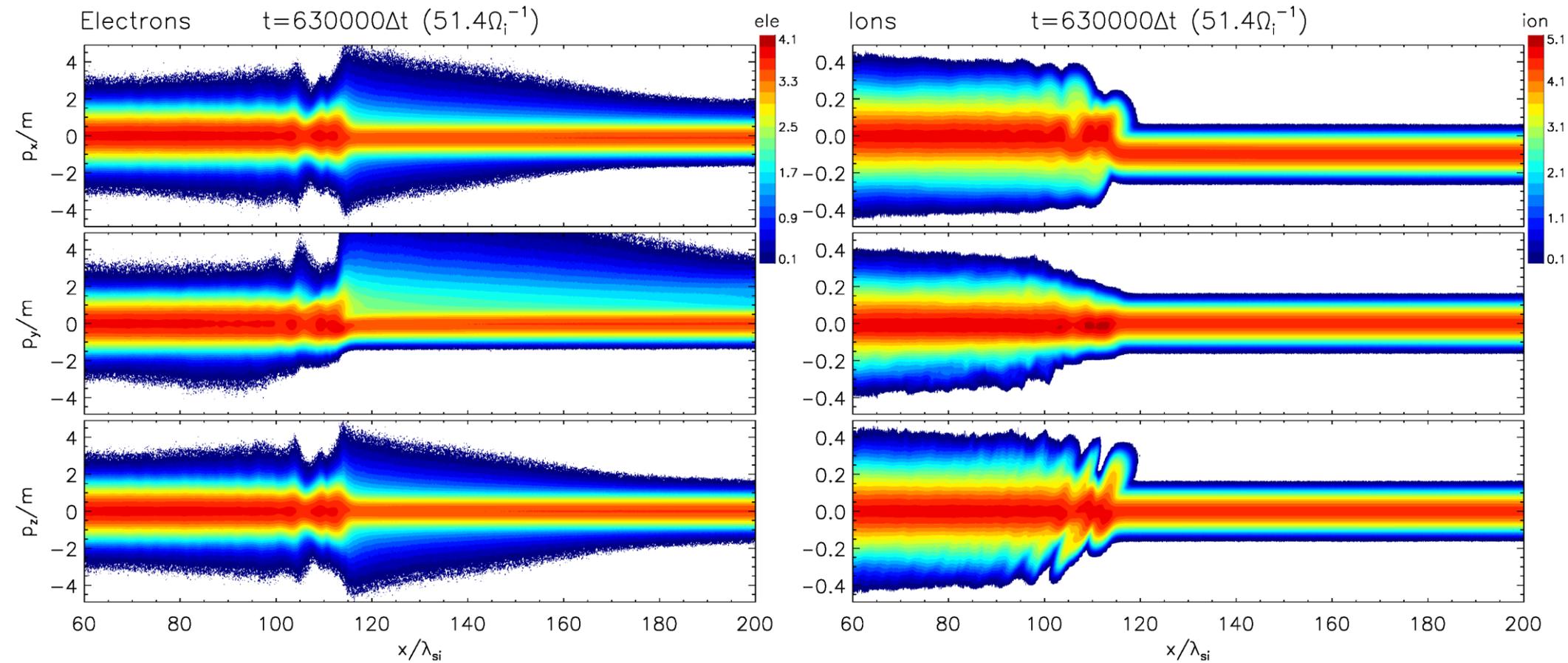


Phase space evolution: *After rippling appearance*

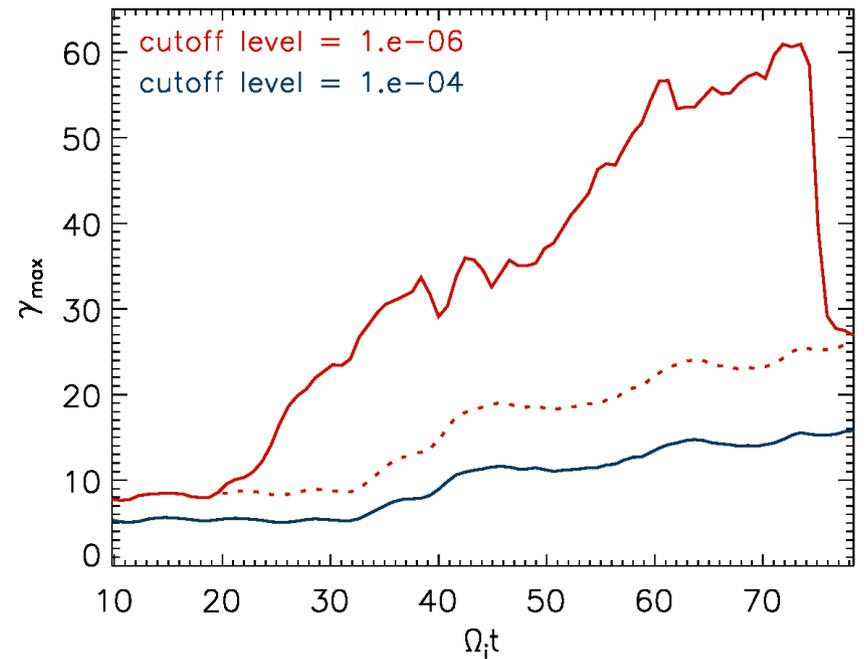
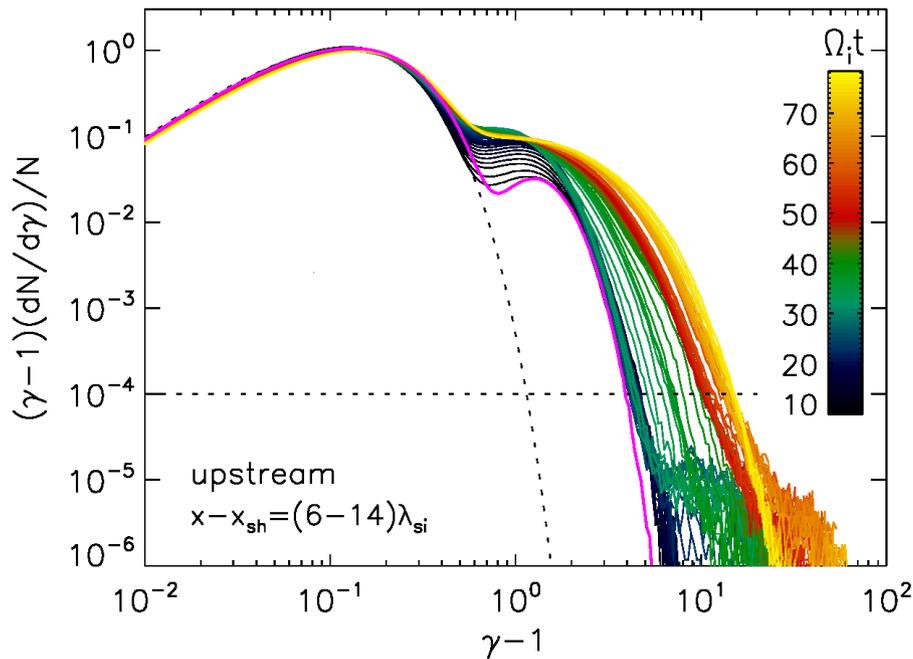
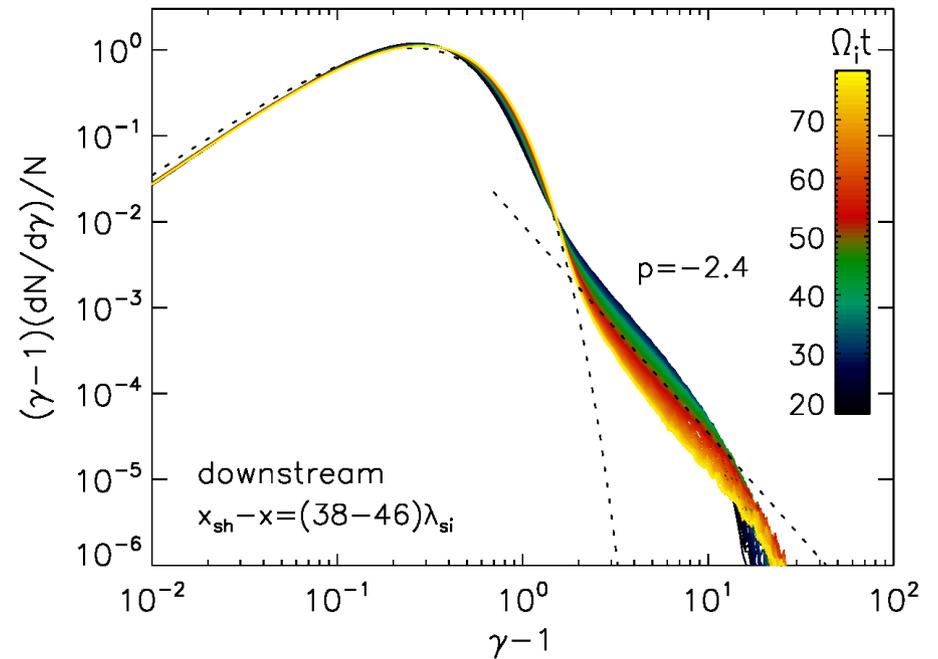
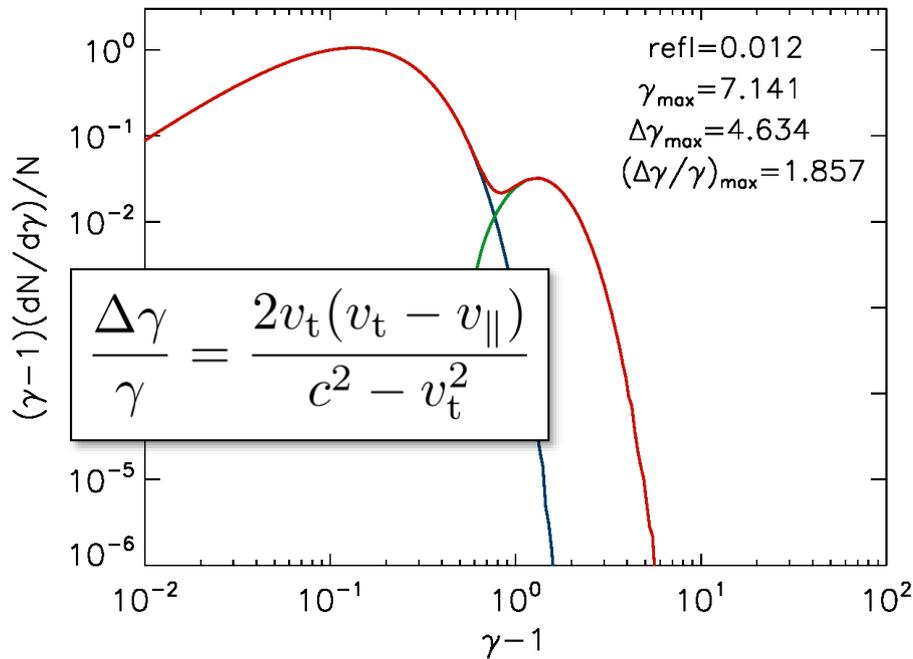


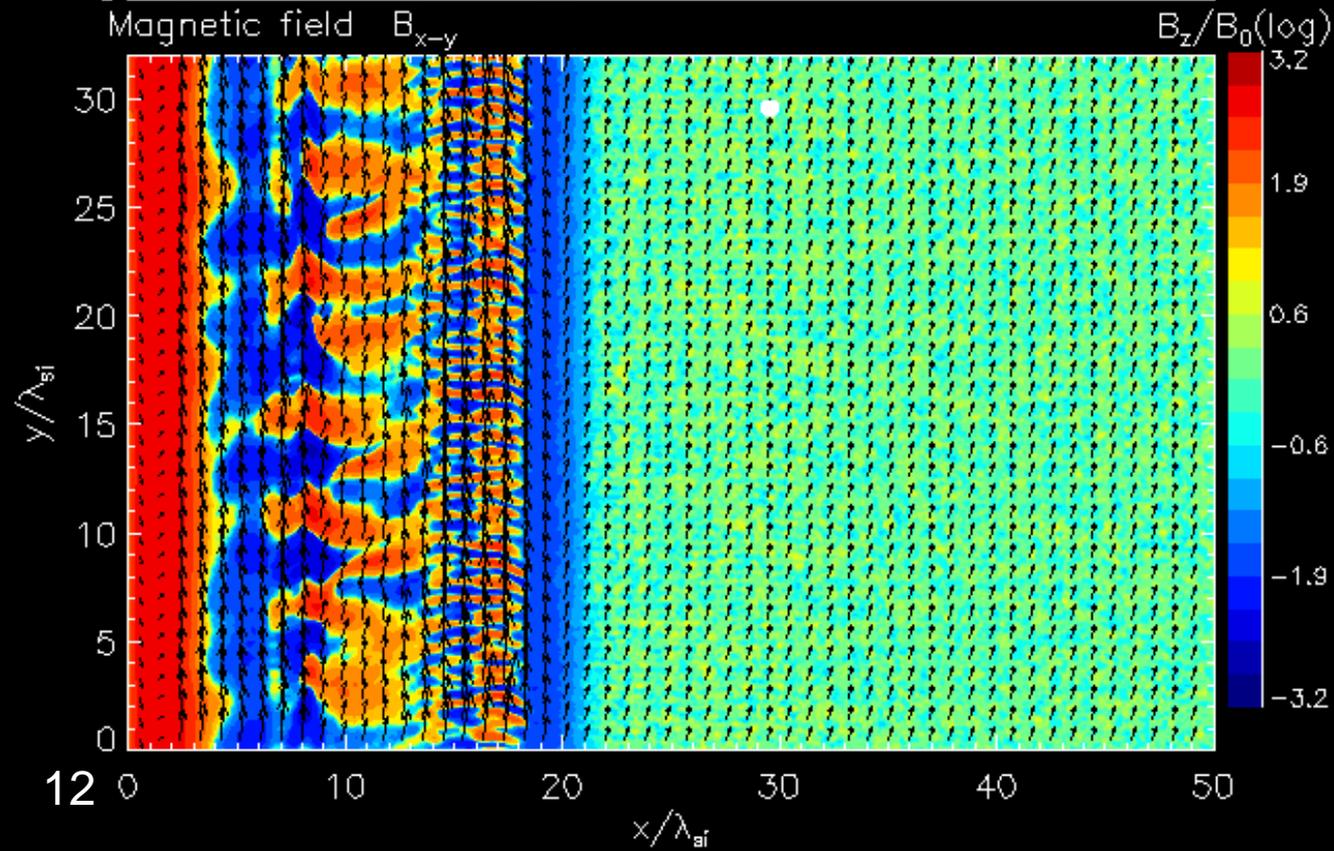
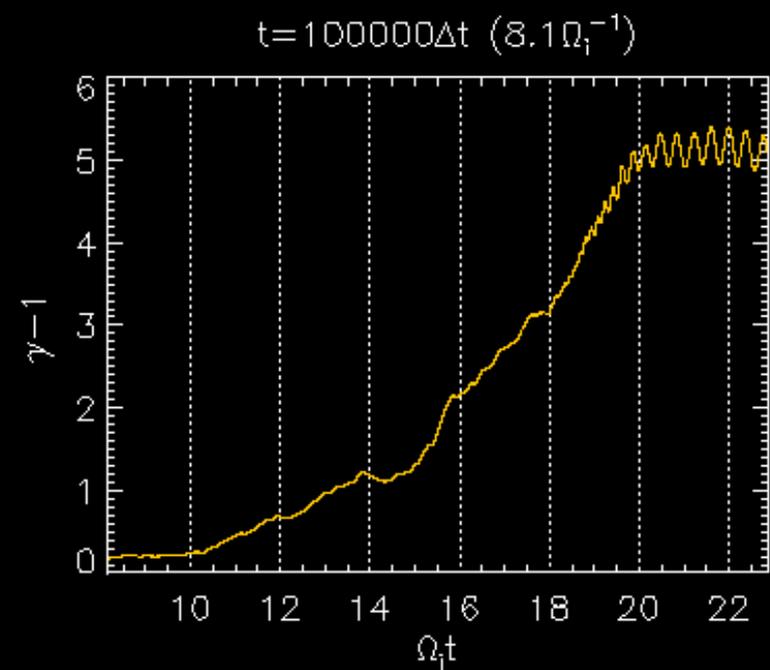
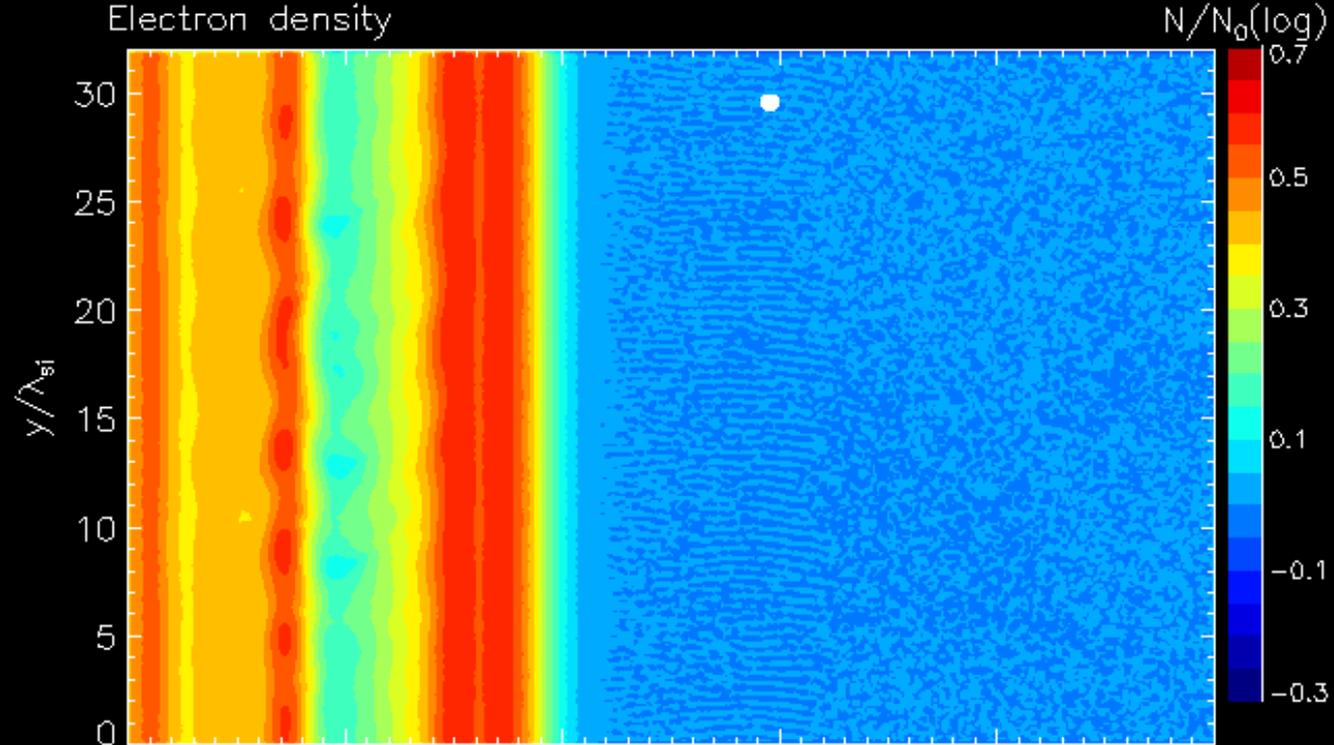
Phase space evolution:

Rippling is well developed

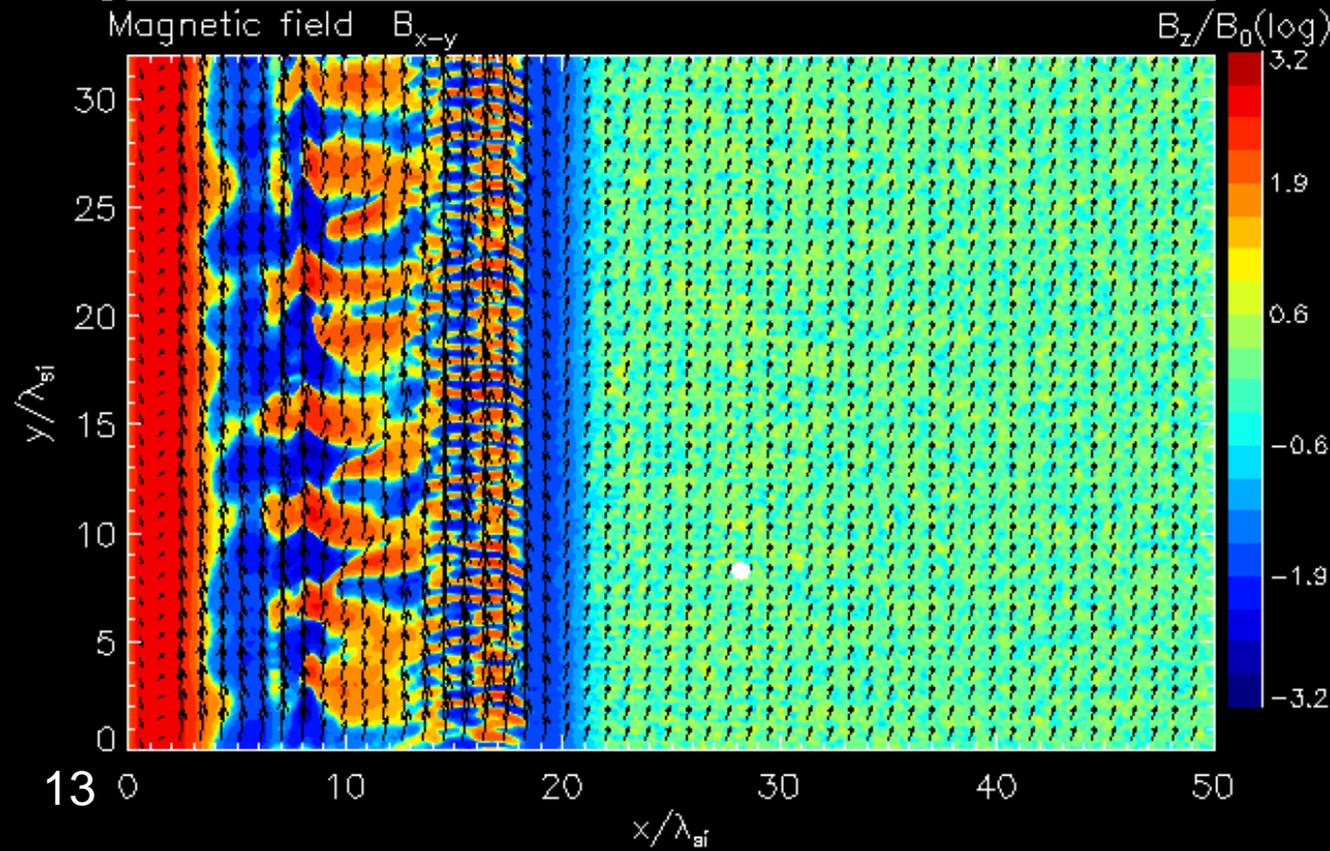
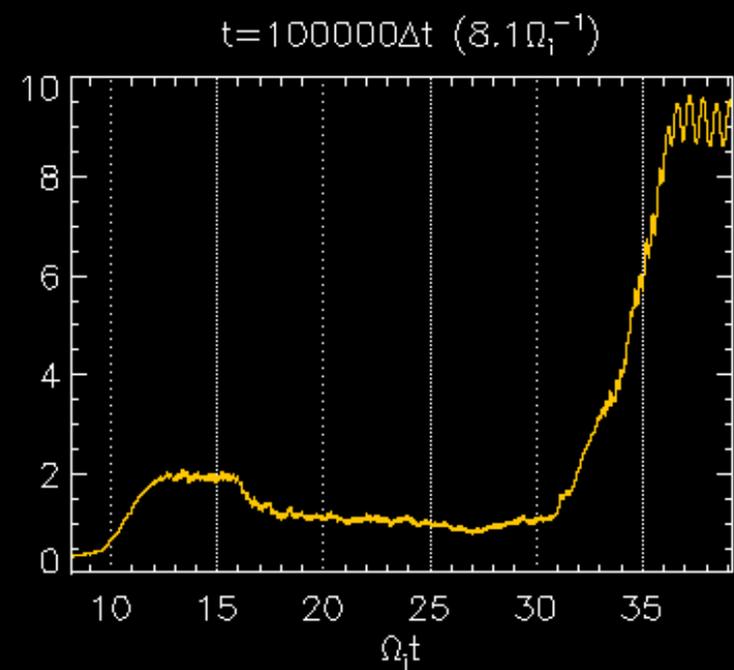
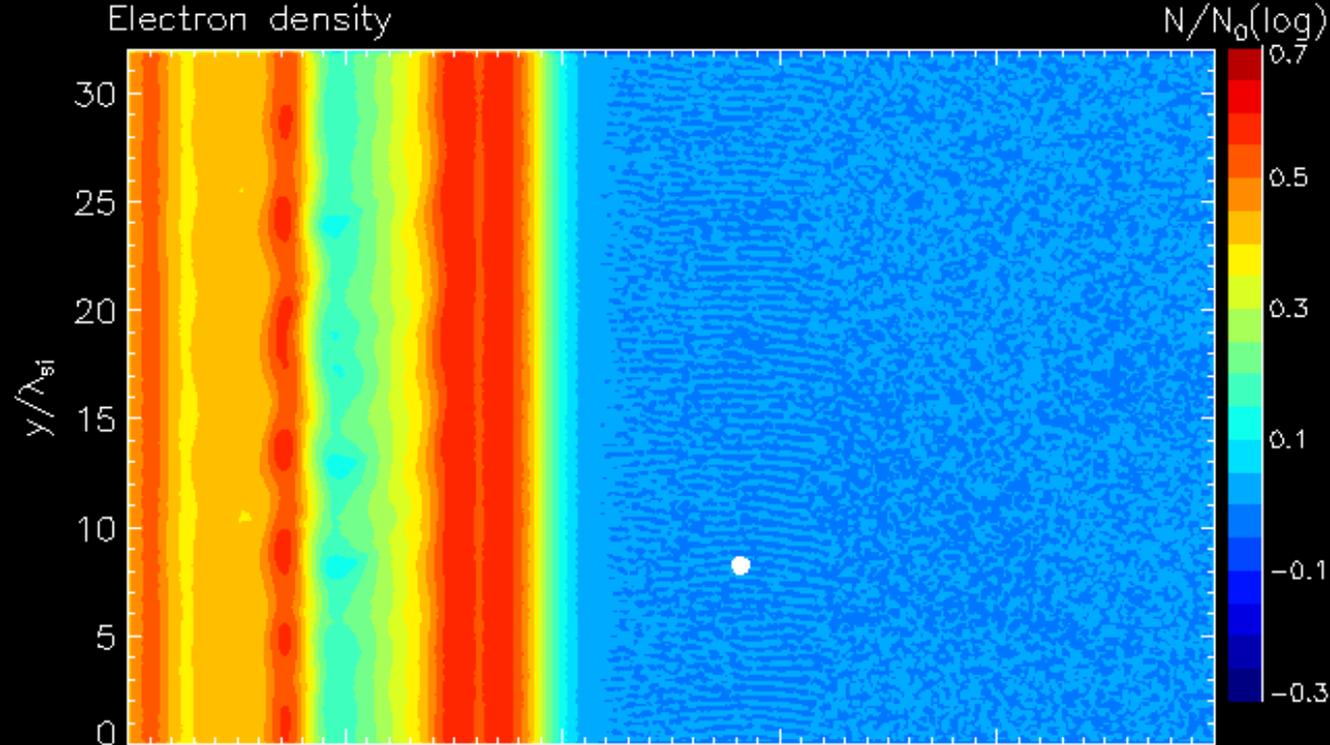


Time evolution of the electron energy spectra

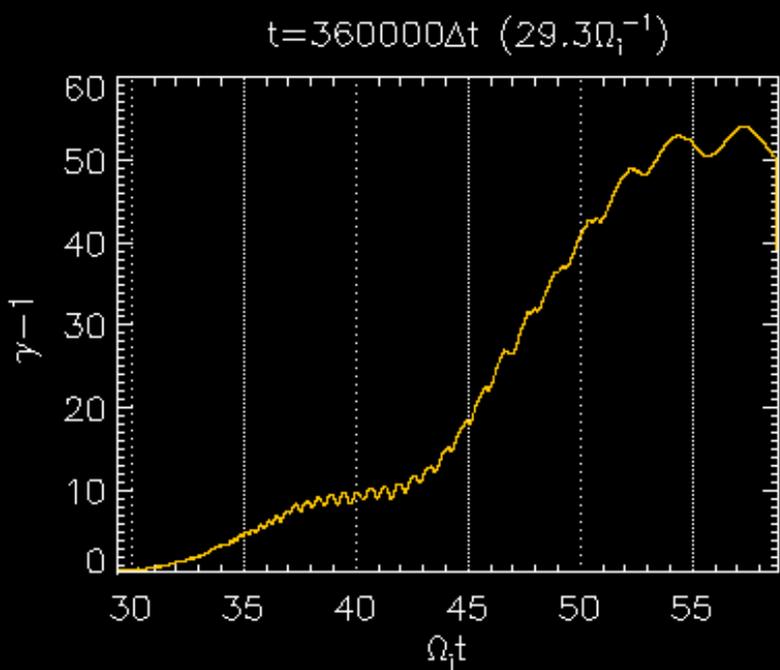
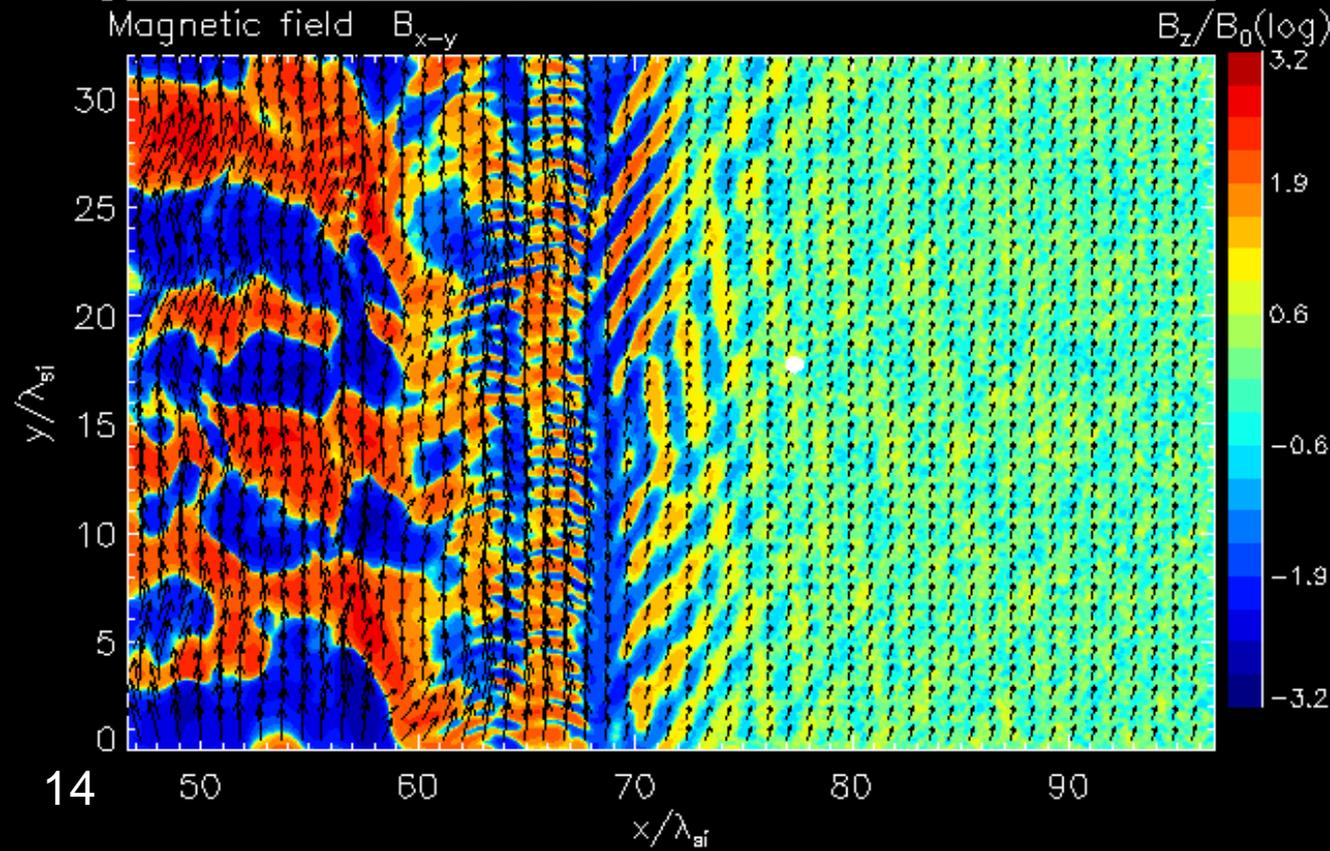
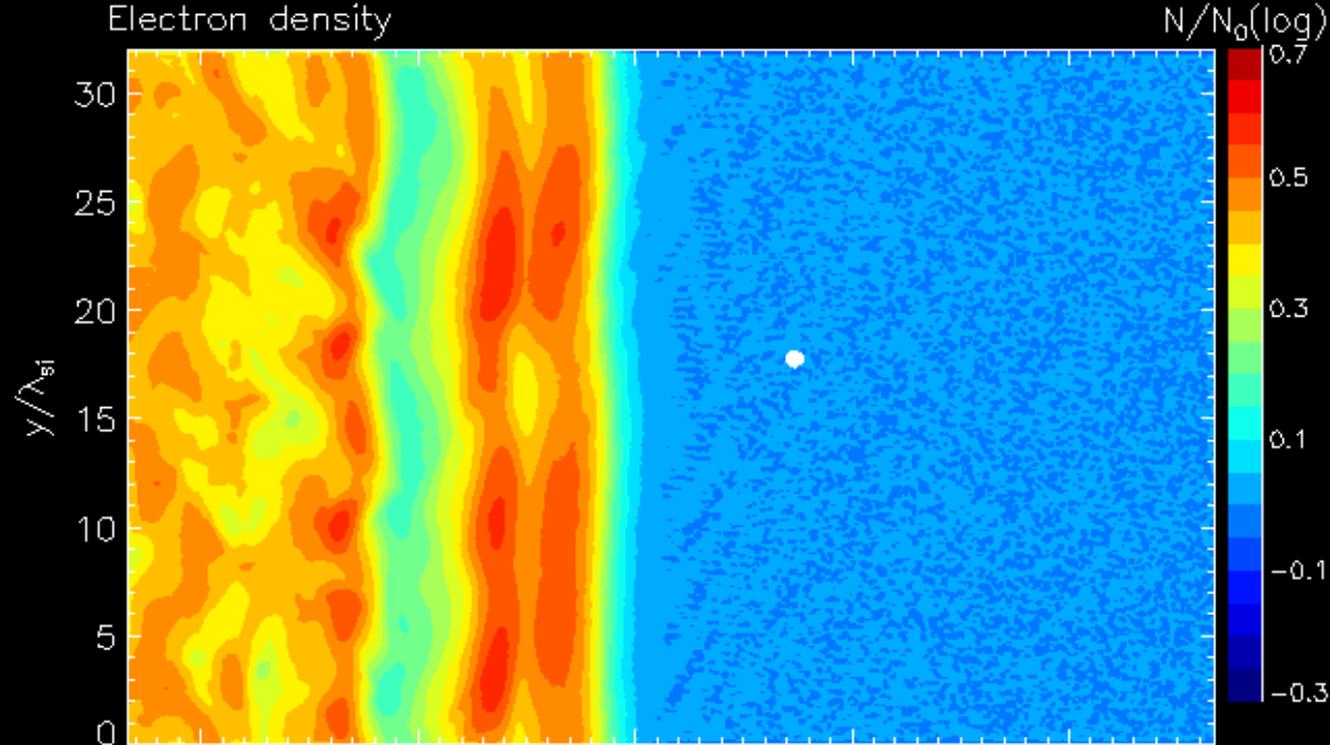




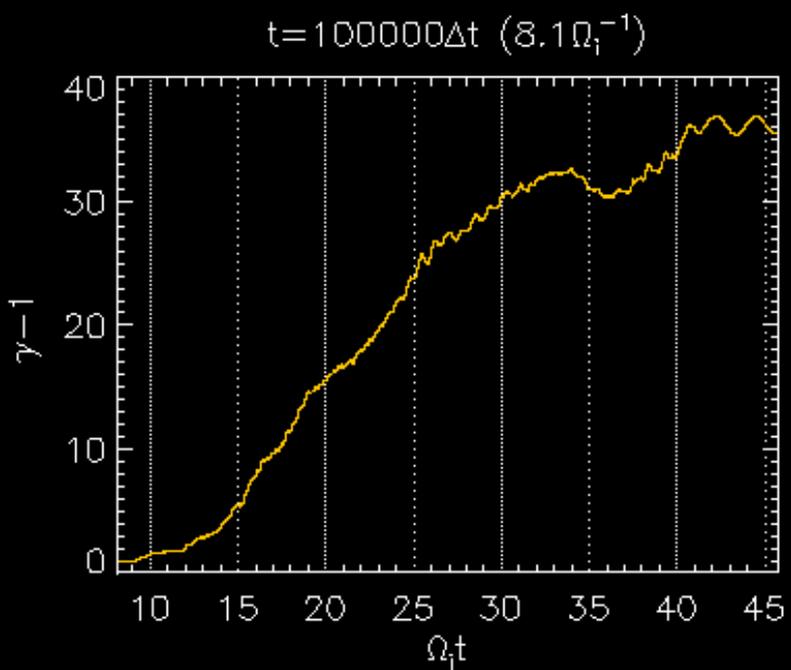
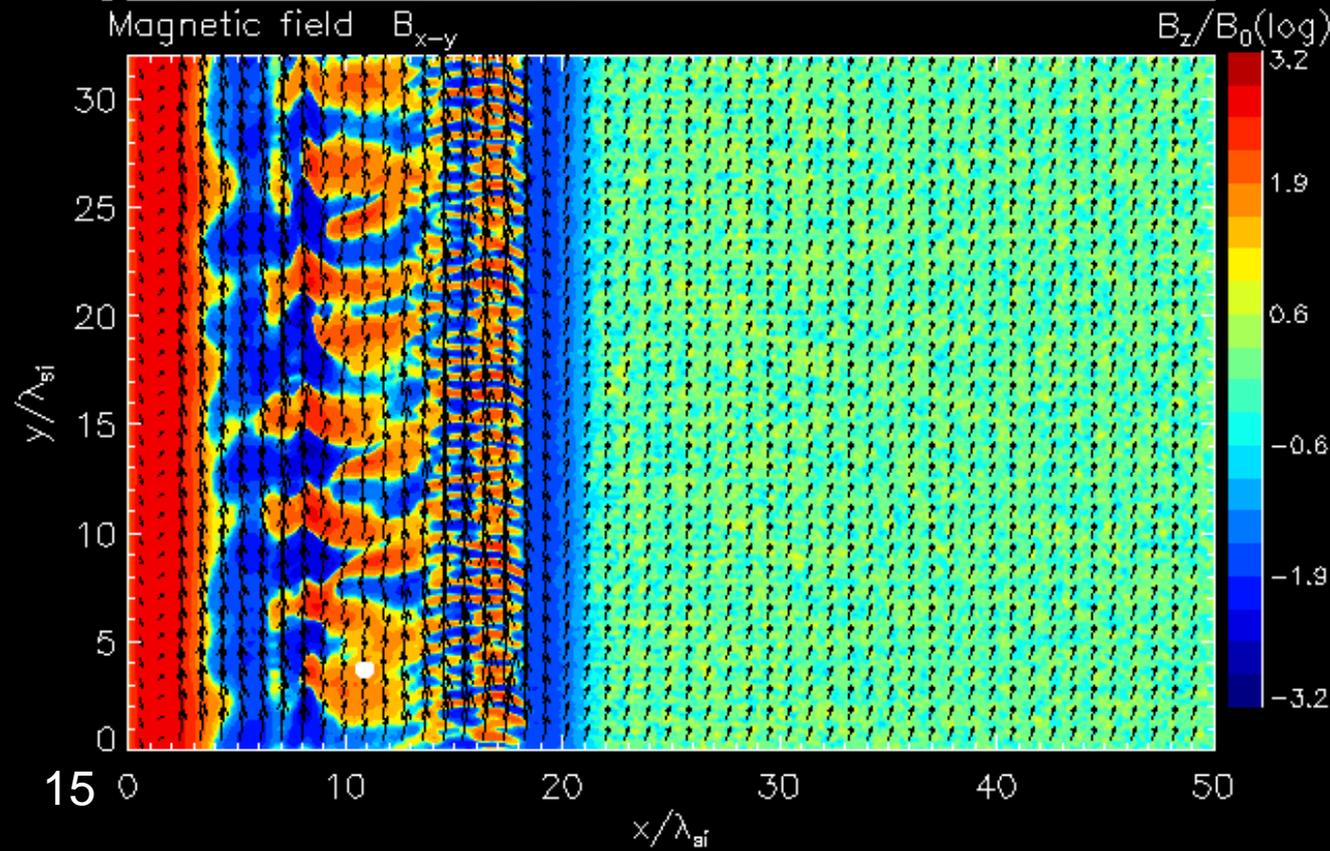
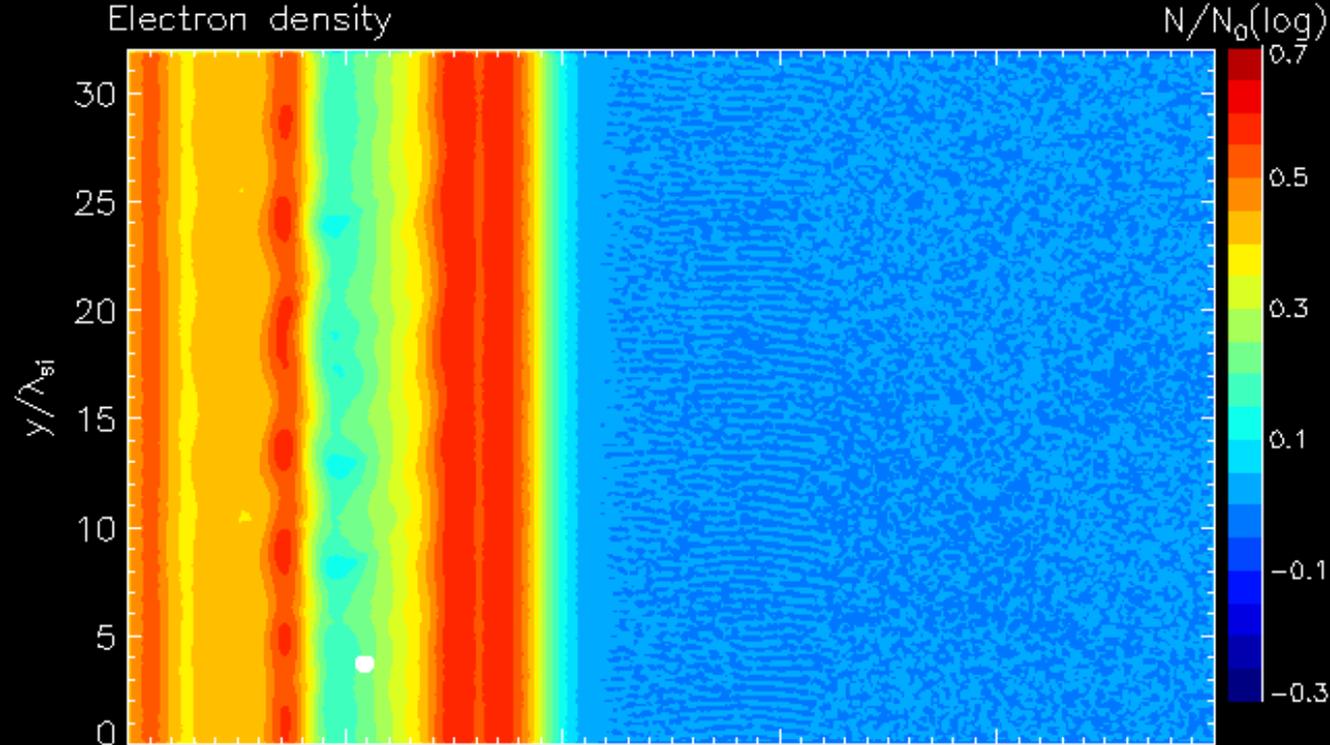
Electron undergoing
the **single-cycle** (regular)
SDA acceleration
process at the shock front



Electron undergoing
the **multi-cycle** (double)
SDA acceleration
process at the shock front

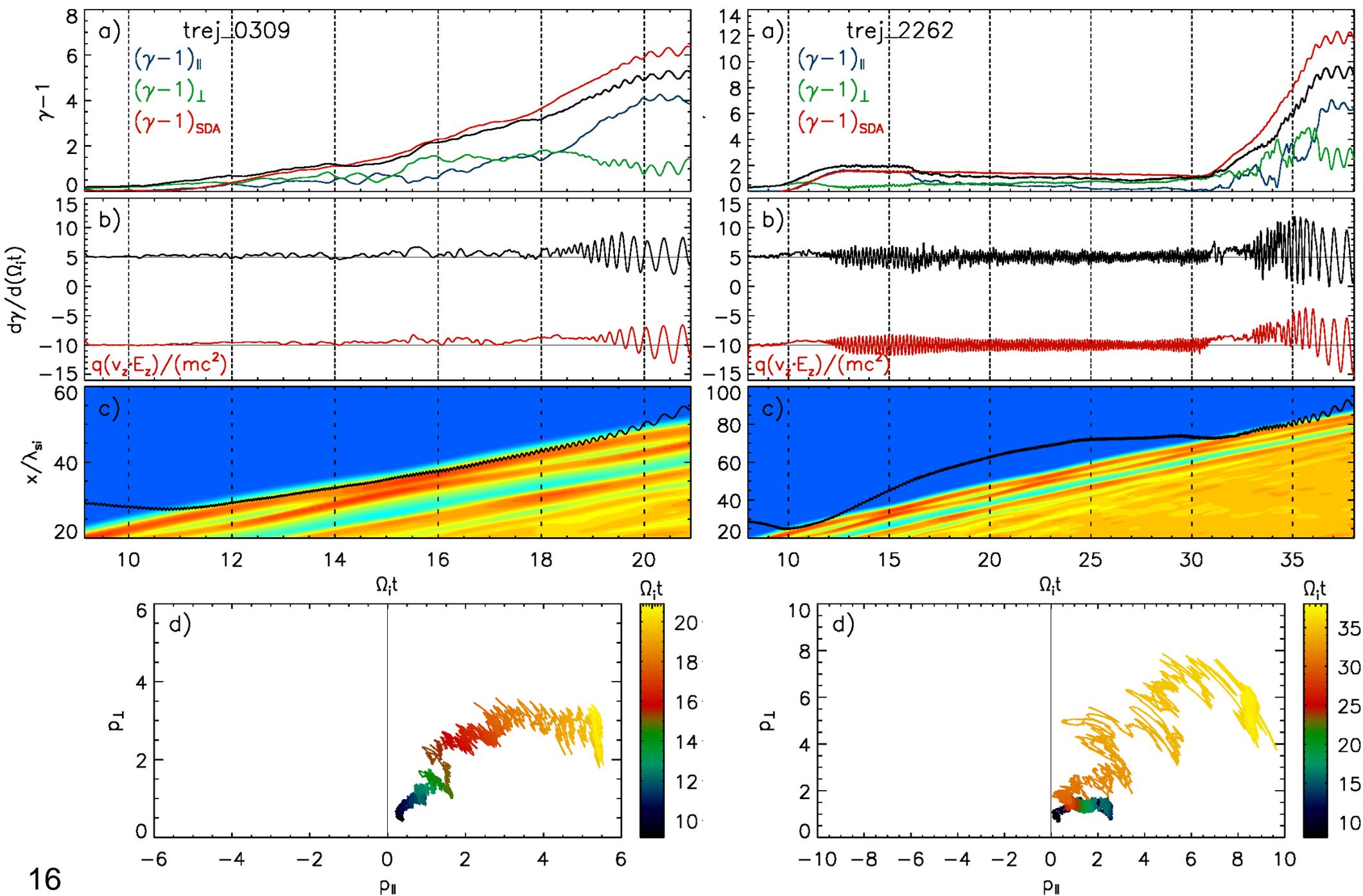


Extra-high energy electron undergoing the **SSDA** acceleration process at front of the **rippled** shock

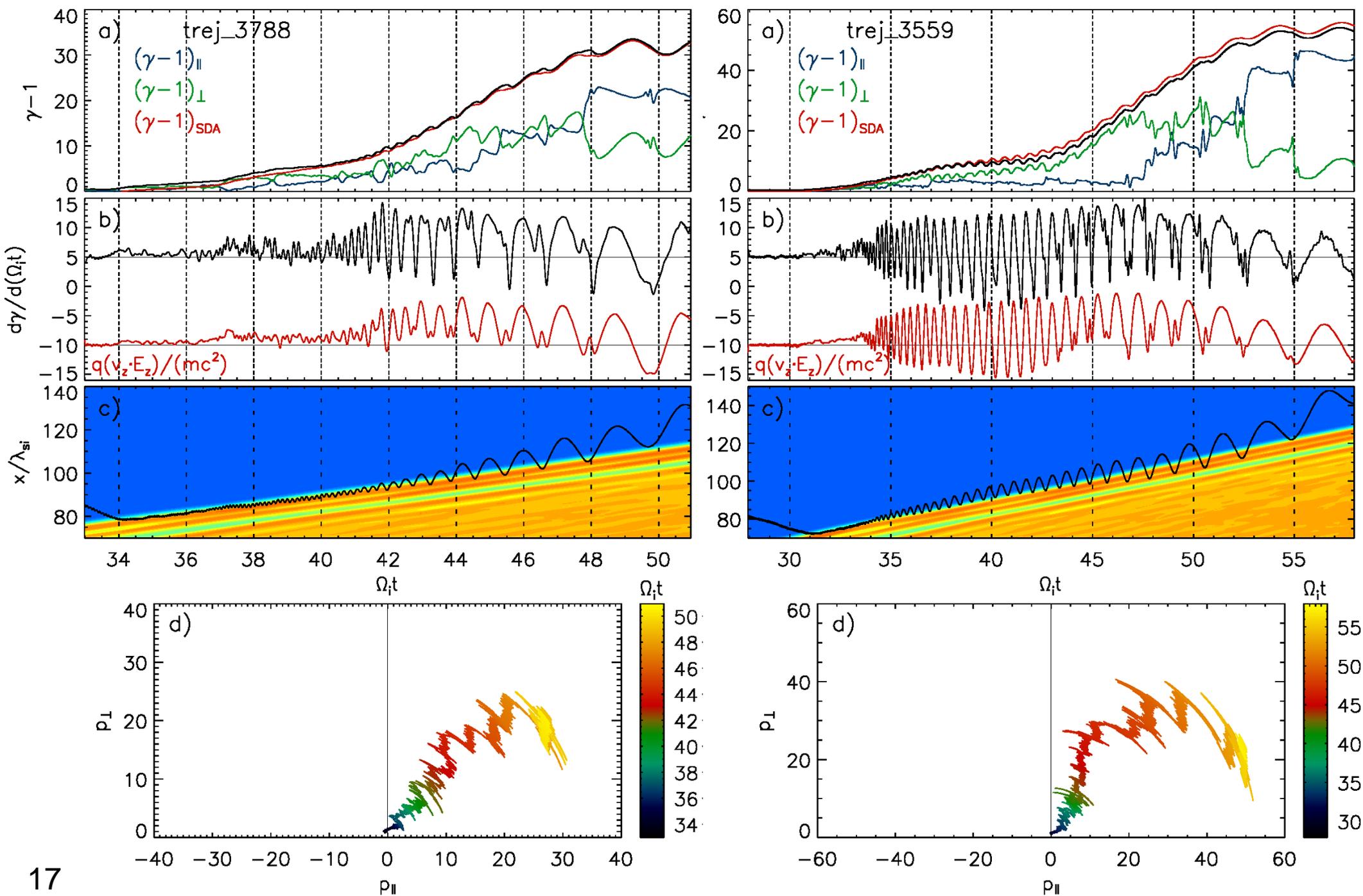


Extra-high energy electron undergoing the **pre-acceleration in the downstream** region before SDA (or SSDA) process at the shock front

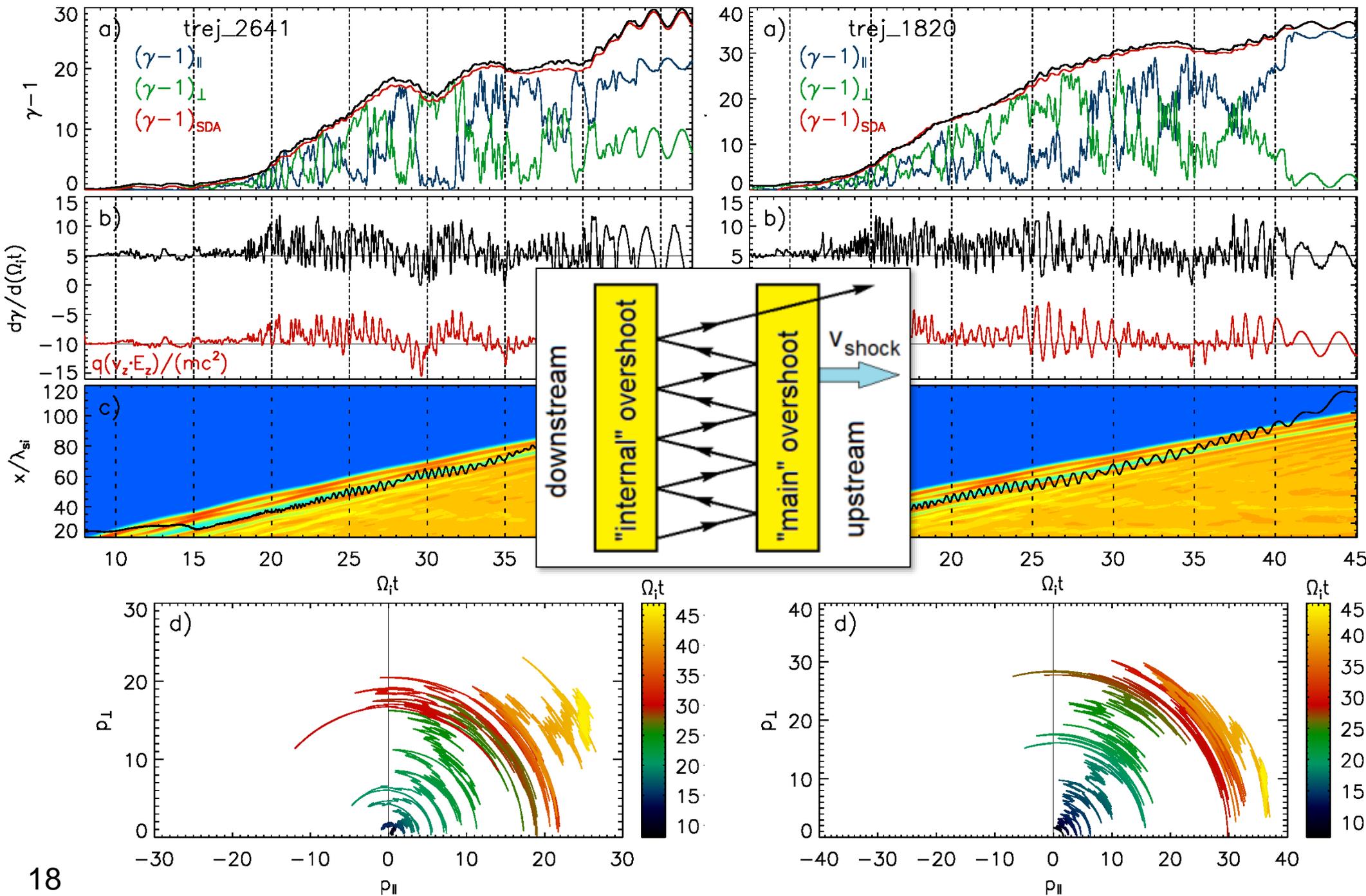
Electron acceleration: SDA



Electron acceleration: *S*SDA



Electron acceleration: *Downstream pre-acceleration*



Summary

- ✓ Performed simulation confirmed all previously described peculiarities of low-Mach-number shocks in high-beta plasma.
- ✓ Acceleration of electrons at laminar shock is found to be in good agreement with SDA theory prediction.
- ✓ Besides of firehose instability, reflected electrons are found to trigger also the bump-on-tail instability, that in turn results in generation of the electrostatic Langmuir waves.
- ✓ Rippling of the shock facilitates the electron acceleration to much higher energies relative to the regular SDA. We interpret this as a result of SSSA process, although rare multi-cycles of SDA are also possible.
- ✓ The population of the most energetic electrons with $\gamma > 50$ are found to be accelerated in the downstream region between two overshoots. We interpret this as a transient effect related with transformation of the shock structure.
- ✓ Downstream electron energy spectra demonstrate non-thermal power-law shape with slope $p = -2.4$, that is consistent with radio-observations.
- ✓ Further work assumes the series of simulations for higher values of plasma beta ($\beta \approx 20$). It requires the usage of several times larger computational resources, up to few tens of thousand CPU-s.